

Prompt Fission Neutron Spectrum Covariances: Impact of Scaling and Normalization

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Representation of Prompt Fission Neutrons in Evaluated Data Libraries

Prompt neutron yield per fission is represented for applications in the following manner (e.g., in the ENDF-6 formats*):

$$Y(E_{ni}, E_{no}) = \nu(E_{ni}) \Psi(E_{ni}, E_{no})$$

$$\int_{0, E_m} \Psi(E_{ni}, E_{no}) dE_{no} = 1 \quad (\text{for all } E_{ni})$$

- E_{ni} = incident neutron energy; E_{no} = outgoing fission neutron energy.
- $\Psi(E_{ni}, E_{no}) \approx 0$ for all $E_{no} > E_m$ ($E_m \leq 20$ MeV**).
- $\nu(E_{ni})$ and $\Psi(E_{ni}, E_{no})$ appear in **separate** evaluated data **files**.
- $\nu(E_{ni})$ and $\Psi(E_{ni}, E_{no})$ are usually **measured** using quite **different experimental techniques**.
- $\nu(E_{ni})$ is usually known to considerably **better accuracy** than $\Psi(E_{ni}, E_{no})$.

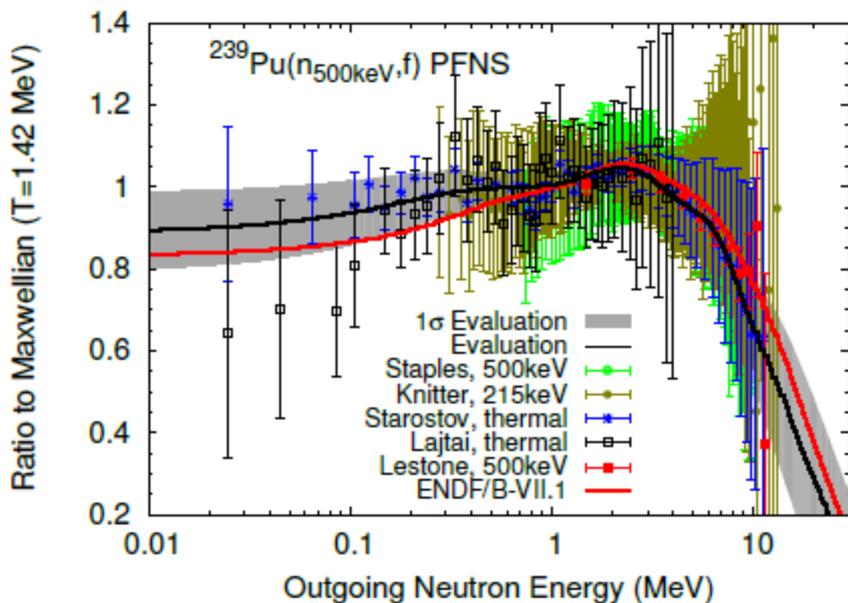
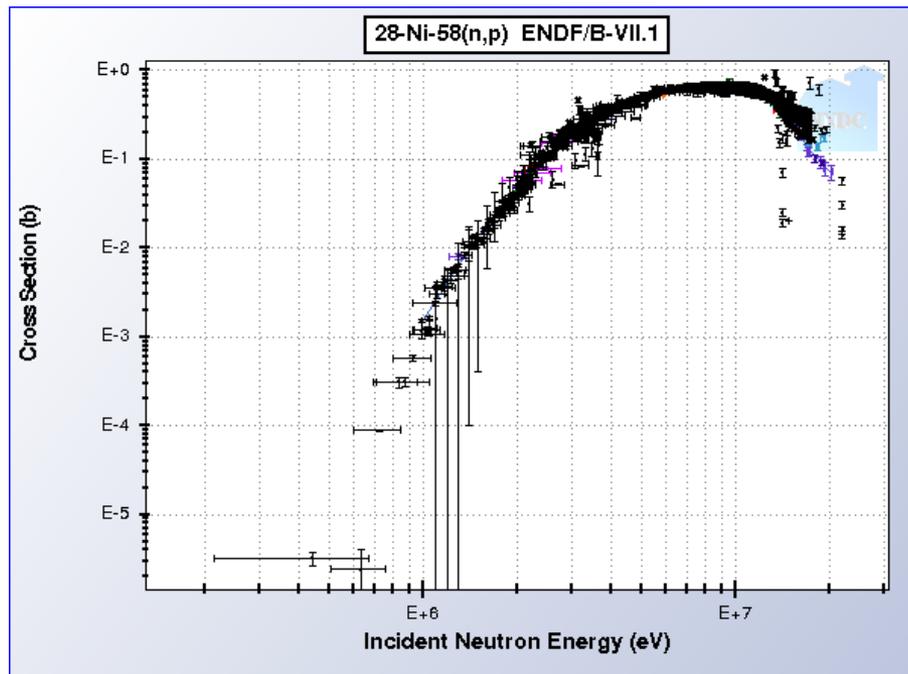
$\Psi(E_{ni}, E_{no})$ is referred to as a **prompt fission neutron spectrum** or, for short, a **PFNS**. It is effectively a **probability distribution** (PDF). $\nu(E_{ni})$ is known as **nu-bar.**, i.e., total number of prompt neutrons per fission.

**ENDF-6 Formats Manual*, ed. M. Herman and A. Trkov, Report BNL-90365-2009, Rev.1 (2010).

** For the major actinides (i.e., $^{235,238}\text{U}$ and ^{239}Pu) nu-bar and PFNS are given in ENDF/B to ≤ 30 MeV, depending on the fissionable isotope. For most applications, 20 MeV is a practical upper limit of emitted neutron energy.

It would appear at first glance that the procedures for evaluating PFNS and cross sections should be essentially the same:

- *Assemble the experimental data.*
- *Adjust these data as needed.*
- *Eliminate poor quality or otherwise questionable values.*
- *Include a model to fill regions poorly represented by experimental data.*
- *Employ least-squares methods.*



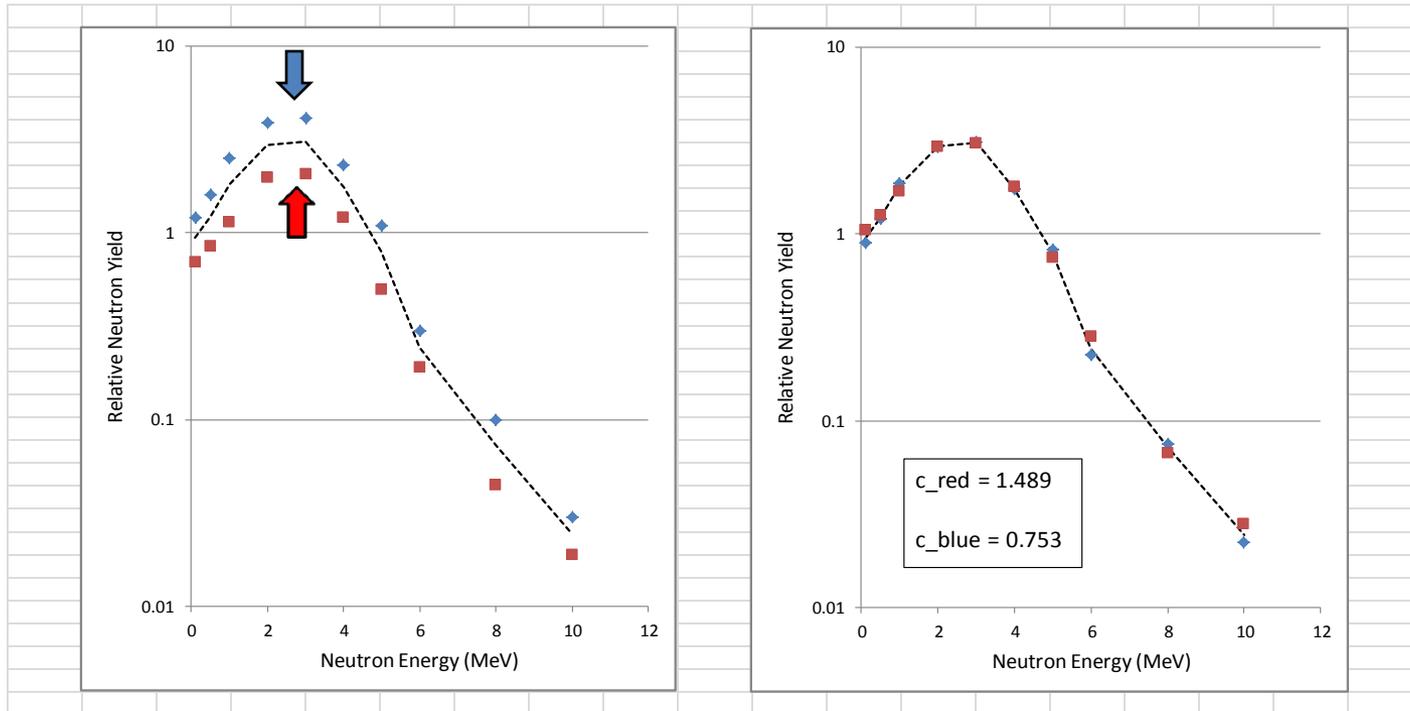
Indeed, there are many **similarities**, but there are certain **differences** that have important implications for applications, especially as related to the **covariance matrices** for the evaluated PFNS as opposed to those for cross sections. These differences stem mainly from:

- *Scaling.*
- *Normalization.*

Scaling? Normalization? What is the Difference?

- These two terms tend to be **used interchangeably** in the literature --- this can lead to **confusion** for PFNS.
- In the case of PFNS, “normalization” is a **particular type** of “scaling” --- there is no change in the “shape” of a spectrum whether it is “scaled” or “normalized”.
- But, there are **profound differences** in the **covariance matrices** for “scaled” and “normalized” PFNS.
- This talk discusses the differences and their impact.

Concept of Scaling



- For a proper evaluation, the **input data** must be **comparable** (no apples vs. bananas).
- Differences in otherwise “comparable” measured PFNS **SHAPE** values arise mainly from different **fission fragment** and/or **neutron detection efficiencies** in the experimental setups.
- Scaled PFNS are PFNS shapes such that each value in a particular set is **multiplied by the same constant** so that the ensemble of included data sets can be treated as “comparable”. Otherwise, **too large chi-square values** will be generated in **least-squares evaluations**.
- There is **no unique way to scale** PFNS. **Optimal scaling minimizes** the solution **chi-square**.
- The **scaling process is relative** ... absolute values of the scaled PFNS are not needed.

Scaling Procedures and Characteristics

Scaling equations (in an energy-group formulation):

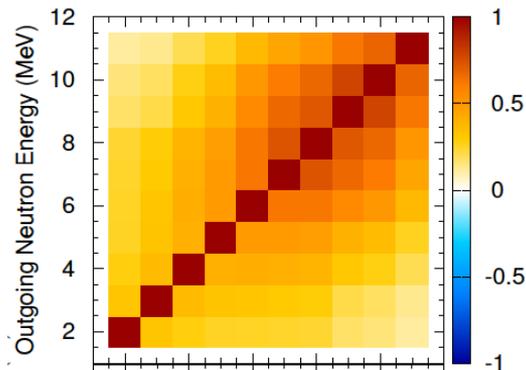
$$\Omega_{ki} = c_k \Phi_{ki} \quad (k=1,K; i=1,n)$$

$$\text{Cov}(\mathbf{\Omega}_k) = (c_k \times c_k) \text{Cov}(\mathbf{\Phi}_k)$$

$\mathbf{\Phi}_k$ is the unscaled PFNS and $\mathbf{\Omega}_k$ is the scaled PFNS.

$\text{Cov}(\mathbf{\Phi}_k)$ is the unscaled covariance matrix and $\text{Cov}(\mathbf{\Omega}_k)$ is the scaled covariance matrix.

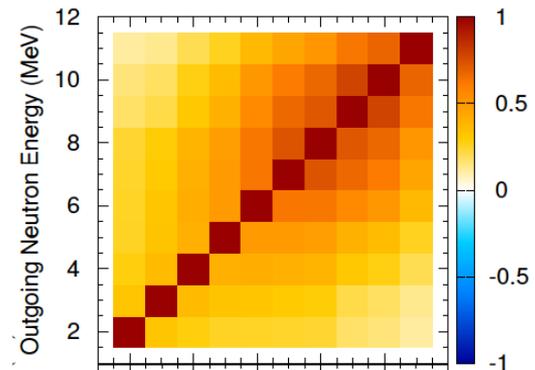
- Scaling **preserves PFNS shape** AND the **correlation pattern** of the **covariance matrix**.



← Before

After →

No change!!



Normalization Procedures and Characteristics

Ω scaled --- but not normalized

$G = \sum_{\text{available}} \Omega_i = \text{arbitrary}$ (collection of available scaled group values for a PFNS Ω).

Ψ normalized

$\sum_{\text{all}} \Psi_i = 1$ (summing every group from zero to an upper energy E_m should yield unity).

Covariance Normalization
Procedure for Group PFNS:

$$\text{Cov}(\Psi) = \mathbf{Q} \times \text{Cov}(\Omega) \times \mathbf{Q}^{\text{transpose}}$$

$$Q_{ij} = (G \cdot \delta_{ij} - \Omega_i) / G^2$$

$\delta_{ij} = 1$ if $i=j$ and $= 0$ otherwise.

The rows and columns of $\text{cov}(\Psi)$ should sum to exactly zero (to within numerical precision of the computational procedure).

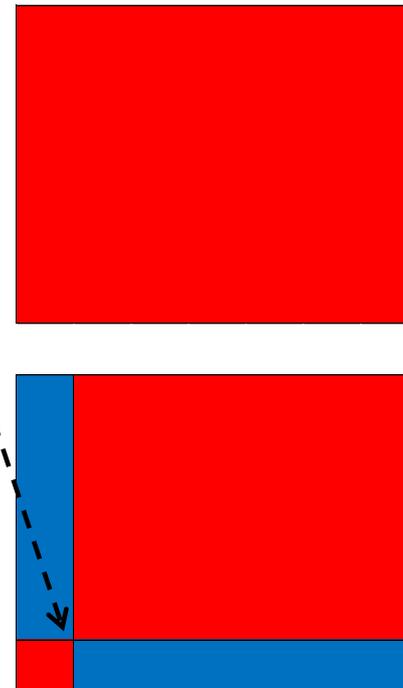
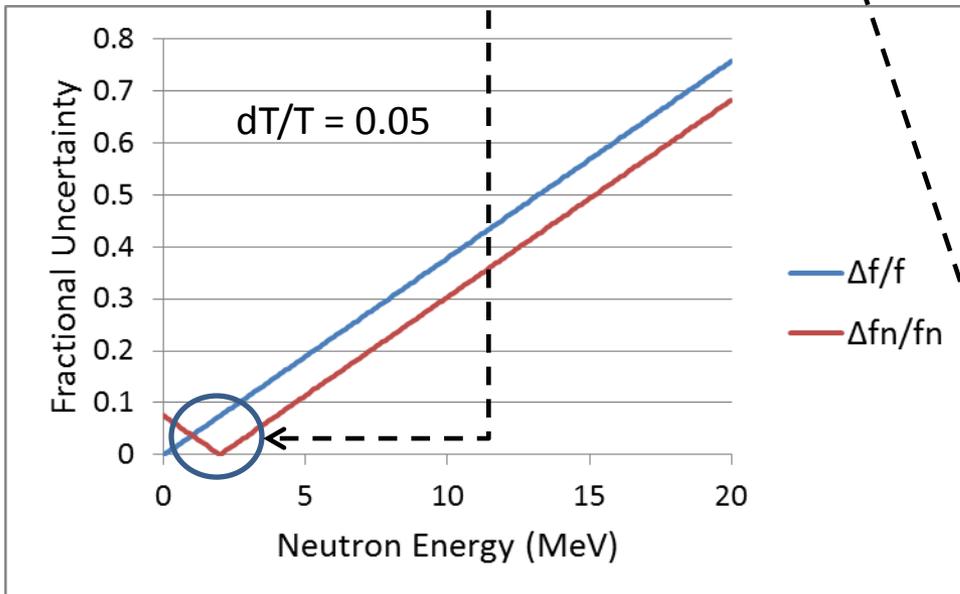
- Any PFNS can be scaled (that's obvious).
- Since a normalized PFNS is a PDF, the present interpretation of "normalization to unity" requires that the representation of a spectrum being normalized should span the entire energy range $(0, E_m)$ where the PFNS values contribute significantly to the energy integral of that spectrum.
- Most experimental PFNS cannot be directly normalized because of their limited energy-range coverage. However, in some instances adequate extrapolations based on models, systematics, or other information can be introduced to enable a largely experimental PFNS to be normalized.
- Most model-generated PFNS are inherently normalized.
- The relationship between a non-normalized and a normalized PFNS is **NON-LINEAR**.
- The matrix transformation from non-normalized $\text{Cov}(\Omega)$ to normalized $\text{Cov}(\Psi)$, as shown here, is linear.

Maxwell-Boltzmann (M-B) Distribution

Unnormalized Formalism		Normalized Formalism
$f = f(E) = C * \text{sqrt}(E) * \exp(-E/T)$		$C_n = [2/\text{sqrt}(\pi)] * [T^{-(3/2)}]$
$df = (df/dT) * dT$	$(T > 0; C = \text{constant})$	$f_n = f_n(E) = C_n * \text{sqrt}(E) * \exp(-E/T)$
$(df/dT) = C * \text{sqrt}(E) * \exp(-E/T) * (E/T/T)$		$(df_n/f_n) = \text{abs}[(E/T) - (3/2)] * (dT/T)$
$(df/f) = (E/T) * (dT/T)$		Pivot Point $\implies (E/T) - (3/2) = 0$

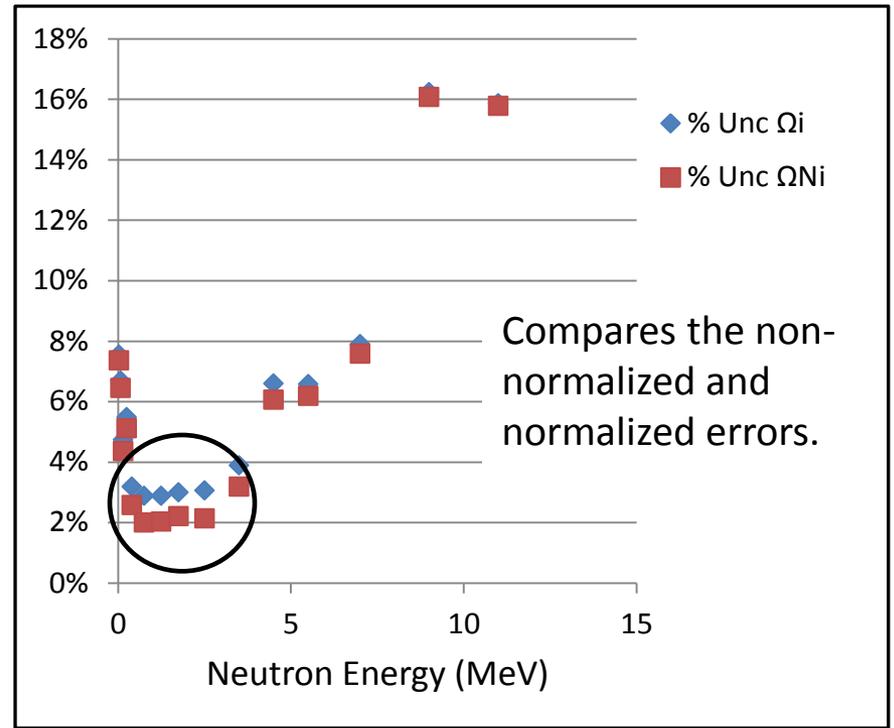
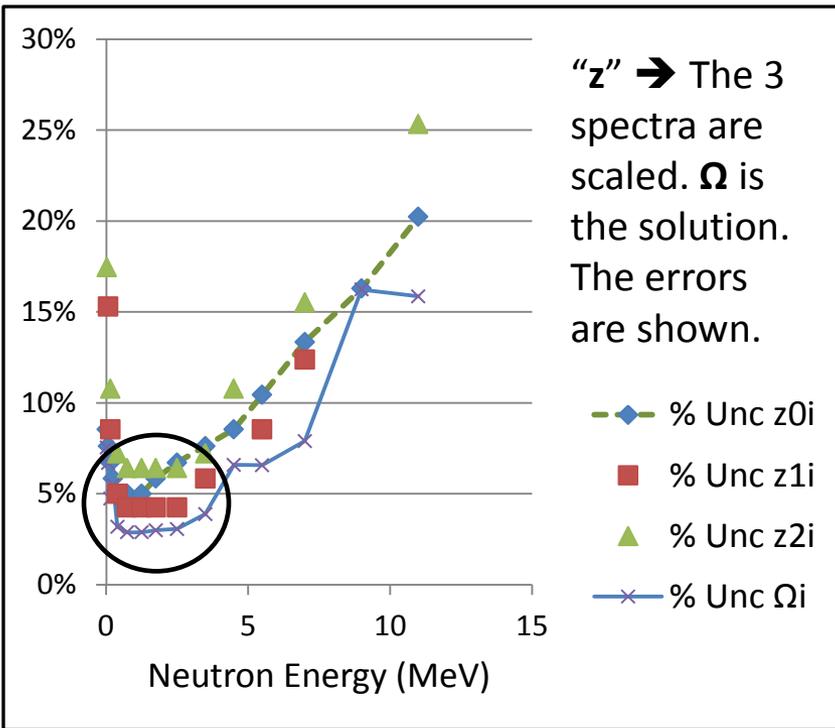
The M-B function is an **extreme case** that demonstrates how normalization affects the errors and correlation matrices for a spectrum shape.

$f(E)$ is non-normalized
 $f_n(E)$ is normalized

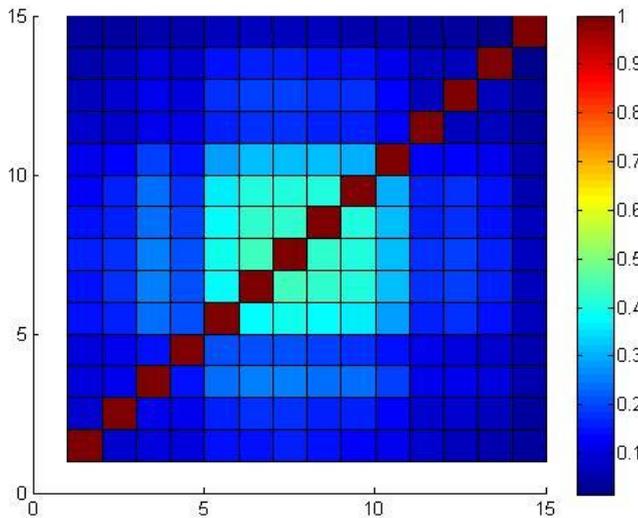


Red \rightarrow
 correlations = +1

Blue \rightarrow
 correlations = -1

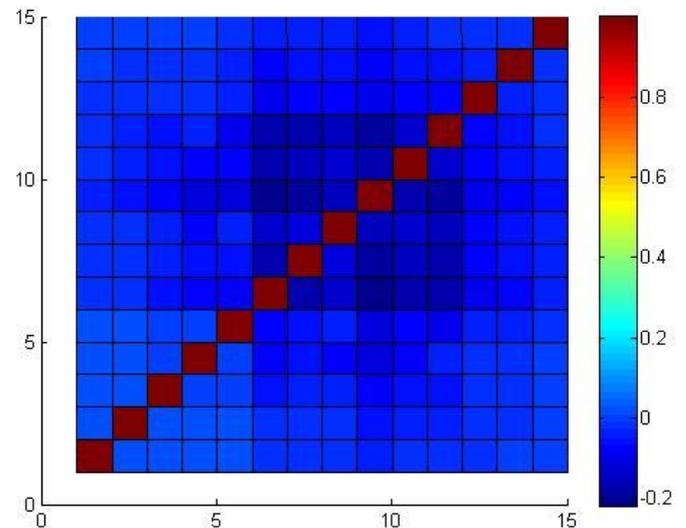


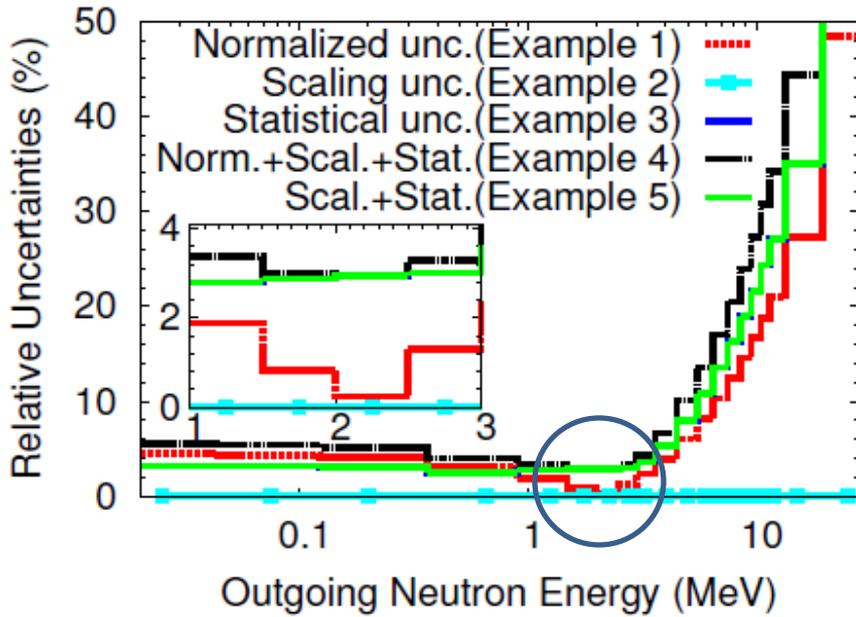
Non-normalized Solution (Ω)



Normalization:
 Reduces error overall, and it also changes the covariance matrix correlation pattern.

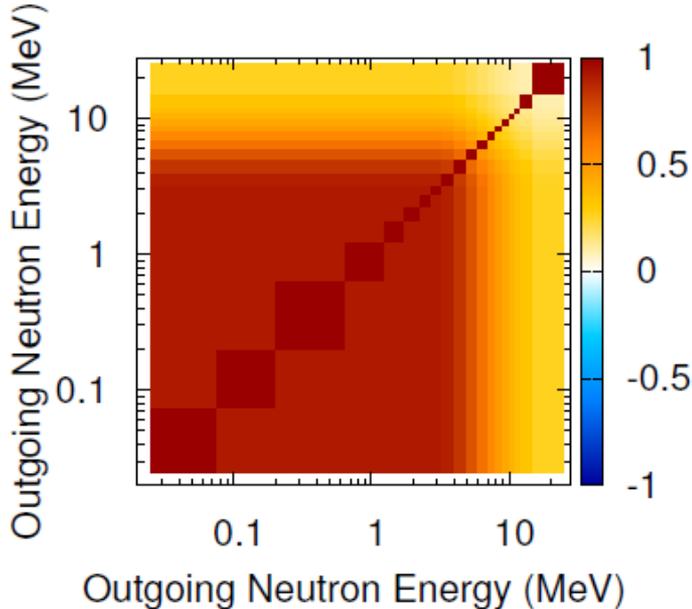
Normalized Solution (Ψ)



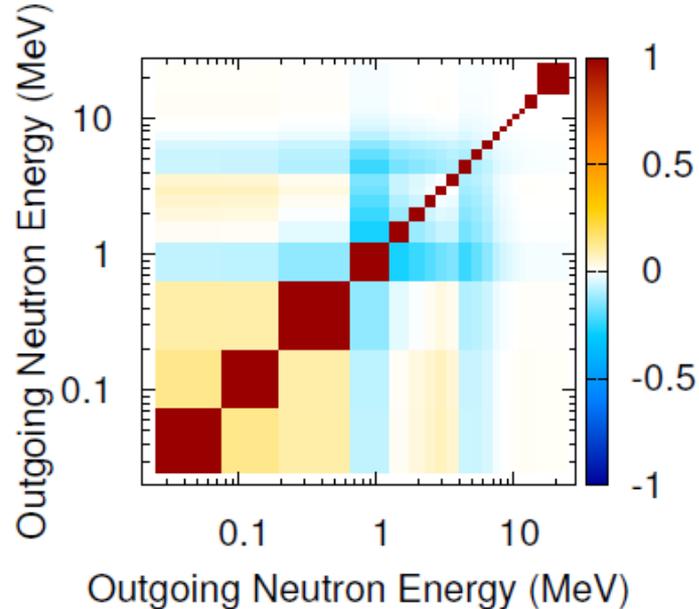


This more **detailed example** illustrates the various **uncertainty components** for the normalized PFNS: Statistical, scaling, and normalization:

- The **scaling uncertainties** (100% correlated across all PFNS data points) are **completely eliminated** upon normalization of a PFNS.
- The **strongly correlated uncertainties** (typical of **model-calculated PFNS**) are reduced by normalization. Unfortunately, it is difficult to conjure PFNS models that do not have relatively strong correlations (**few parameters**).
- So, **inclusion of model PFNS data** can lead to evaluated PFNS with **too small errors**.



(a) Before normalization.



(b) After normalization

Notice the dramatic difference in correlation patterns!!

Key Points and Some References

- Collections of PFNS need to be scaled so as to be “comparable” prior to their evaluation.
- Scaling is a “relative” process and not an “absolute” one.
- PFNS scaling does not affect the covariance matrix correlations.
- Optimal scaling leads to the minimal chi-square possible for the evaluated solution.
- “Complete” PFNS that span the whole spectrum energy range can be normalized.
- PFNS normalization to unity (a probability distribution) is a non-linear process.
- Normalization does not change the PFNS shape but it has a dramatic effect on the properties of the covariance matrix.
- Scaling uncertainties (100% correlated across all PFNS data points) vanish when the spectrum is normalized.
- Model PFNS tend to be inherently normalized (e.g., the Los Alamos Model).
- The strong correlations of models tend to lead to unrealistically small evaluated PFNS uncertainties, especially near the “pivot point”, when the spectrum is normalized.
- It is generally best to evaluate non-normalized (scaled) PFNS and then normalize the evaluated solution afterwards, as required to satisfy ENDF-6 format requirements.
- However, mixed non-normalized and normalized PFNS can be evaluated if properly scaled.
- The expected minimal uncertainty (pivot point) for an evaluated PFNS should be defined by the experimental “shape” uncertainty, i.e., the uncertainty AFTER normalization ($\approx 1\%$).

*D. Neudecker et al., *Impact of the Normalization Condition and Model Information on Evaluated Prompt Fission Neutron Spectra and Associated Uncertainties*, to be published in NSE (2015).

*D. Neudecker et al., *Evaluation of the ^{239}Pu Prompt Fission Neutron Spectrum Induced by Neutrons of 500 keV and Associated Covariances*, submitted to NIM-A for publication (2015).

*D. Smith et al., *Prompt Fission Neutron Spectrum Evaluation Techniques*, Report INDC(NDS)-0678 (2015).